## Chapter 2: Digital Image Fundamentals Structure of the Human Eye

The iris acts as a diaphragm to control the amount of light entering the eye.

Three membranes enclose the eye: Cornea and sclera, Choroid, Retina


Pupil size: 2-8mm
Eye color: melanin (pigment) in iris



The distribution of rods and cones is radially symmetric wrt the fovea (central portion of the retina), except at the blind spot which includes no receptors.

Cones are responsible for photopic (color or bright-light) vision; while rods are for scotopic (dim-light) vision.

Fovea area in the retina is circular with 1.5 mm in diameter where most of the cones are concentrated with 150000 cones $/ \mathrm{mm}^{2}$. This is easily achievable with medium ${ }^{2.5}$ resolution CCD imasing chin of size $5 \mathrm{~mm} \times 5 \mathrm{~mm}$ !

How's an object seen at the back of the eye?


The focal lenght (distance bet center of the lens and the retina) varies from 17 mm to 14 mm (as the refractive power of the lens increases from its minimum to its maximum). Recall that $H / L=h / l$

Perception takes place by the relative excitation of light receptors, which transform radiant energy into electrical impulses that are ultimately decoded by the brain.

Human eye can adapt to an enormous range (in the order of $10^{10}$ ) of light intensity levels, from scotopic threshold to the glare limit.

Subjective brightness (i.e. perceived intensity) is a logarithmic function of the light intensity incident on the eye.


In photopic vision alone, the range is about $10^{6}(-2$ to 4 in the $\log$ scale $)$.
The transition from scotopic to photopic vision is gradual over the range $(0.001,0.1)$ millilambert ${ }^{1}(-3$ to -1 mL in the $\log$ scale $) .{ }^{2}$

- Scotopic means vision under low light condition
- Ptotopic means under normal/high light condition
- HVS cannot operate over $10^{\wedge} 10$ orders of magnitude dynamic range simultaneously
- Total range of distinct intensity levels eye can discriminate simultaneously is small
- For given conditions, the current visual sensitivity of HVS is called "brightness adaptation".
incremental illumination $\Delta I$ appears in the form of a short duration flash


FIGURE 2.5 Basic
experimental setup used to characterize brightness discrimination.

Note: $I$ is luminance, measured in $\mathrm{cd} / \mathrm{m}^{2}$


FIGURE 2.6
Typical Weber ratio as a function of intensity.

A small Weber ratio indicates "good" brightness where a small percentage change in illumination is discriminable. On the other hand, a large Weber ratio represents "poor" brightness indicating that a large percentage change in intensity is needed.

The curve shows that brightness discrimination is poor (large Weber ratio) at low level of illumination, and it improves significantly (Weber ratio decreases) as background illumination increases.

The two branches illustrate the fact that at low levels of illumination, vision is carried out by the rods, whereas at high levels (showing better discrimination), cones are at work.

2.11


| a |
| :--- |
| b |
| c |

FIGURE 2.7
Illustration of the
Mach band effect.
Perceived
intensity is not a simple function of actual intensity.

a b c
FIGURE 2.8 Examples of simultaneous contrast. All the inner squares have the same intensity, but they appear progressively darker as the background becomes lighter.


Energy of one photon (electron volts)


FIGURE 2.10 The electromagnetic spectrum. The visible spectrum is shown zoomed to facilitate explanation, but note that the visible spectrum is a rather narrow portion of the EM spectrum.

The electromagnetic spectrum can be expressed in terms of wavelength $(\lambda)$, frequency $(v)$, or energy $(E)$. Recall that

$$
\lambda=c / v
$$

where $c$ is the speed of light $\left(2.998 \times 10^{8} \mathrm{~m} / \mathrm{s}\right)$.
The energy of the various components is given by:

$$
E=h v
$$

where $h$ is Planck's constant ( $6.62606891 \times 10^{-34}$ Jouleseconds (or $\mathrm{m}^{2} \mathrm{~kg} / \mathrm{s}$ )). $E$ is measured in electron-volt.

## Light and EM Spectrum

The colors that humans perceive in an object are determined by the nature of the light reflected from the object.
e.g. green objects reflect light with wavelengths primarily in the 500 to 570 nm range while absorbing most of the energy at other wavelength

## Monochromatic light: void of color

Intensity is the only attribute, from black to white Monochromatic images are referred to as gray-scale images

Chromatic light bands: 0.43 to 0.79 um
The quality of a chromatic light source:
Radiance: total amount of energy
Luminance (Im): the amount of energy an observer perceives from a light source
Brightness: a subjective descriptor of light perception that is impossible to measure. It embodies the achromatic notion of intensity and one of the key factors in describing color sensation.

FIGURE 2.11
Graphical
representation of one wavelength.


(a) Single imaging
sensor.
(b) Line sensor.
(c) Array sensor.
a
b
FIGURE 2.12
,
c

## Transform

 illumination energy into digital images


FIGURE 2.13
Combining a single sensor with motion to generate a 2-D image.

a b
FIGURE 2.14 (a) Image acquisition using a linear sensor strip. (b) Image acquisition using a circular sensor strip.


FIGURE 2.15 An example of the digital image acquisition process. (a) Energy ("illumination") source. (b) An element of a scene. (c) Imaging system. (d) Projection of the scene onto the image plane. (e) Digitized image.

## Image formation model

$f(x, y)=i(x, y) r(x, y)$
$f(x, y)$ : intensity at the point $(x, y)$
$i(x, y)$ : illumination at the point $(x, y)$
(the amount of source illumination incident on the scene) $r(x, y)$ : reflectance/transmissivity at the point $(x, y)$ (the amount of illumination reflected/transmitted by the object) where $0<i(x, y)<\infty$ and $0<r(x, y)<1$

## Some Typical Ranges of illumination

- Illumination

Lumen - A unit of light flow or luminous flux
Lumen per square meter $\left(\mathrm{lm} / \mathrm{m}^{2}\right)$ - The metric unit of measure for illuminance of a surface

- On a clear day, the sun may produce in excess of $90,000 \mathrm{~lm} / \mathrm{m}^{2}$ of illumination on the surface of the Earth
- On a cloudy day, the sun may produce less than $10,000 \mathrm{~lm} / \mathrm{m}^{2}$ of illumination on the surface of the Earth
- On a clear evening, the moon yields about $0.1 \mathrm{~lm} / \mathrm{m}^{2}$ of illumination
- The typical illumination level in a commercial office is about $1000 \mathrm{~lm} / \mathrm{m}^{2}$


## Some Typical Ranges of Reflectance

## - Reflectance

- 0.01 for black velvet
- 0.65 for stainless steel
- 0.80 for flat-white wall paint
- 0.90 for silver-plated metal
- 0.93 for snow


a b
FIGURE 2.17 (a) Continuous image projected onto a sensor array. (b) Result of image sampling and quantization.

a
b
c

FIGURE 2.18
(a) Image plotted
as a surface.
(b) Image displayed as a visual intensity array.
(c) Image shown as a 2-D numerical array
( $0, .5$, and 1 represent black, gray, and white, respectively).


FIGURE 2.19 An
image exhibiting
saturation and noise. Saturation is the highest value beyond which all intensity levels are clipped (note how the entire
saturated area has a high, constant intensity level).
Noise in this case
appears as a grainy texture pattern. Noise, especially in the darker regions of an image (e.g., the stem of the rose) masks the lowest detectable true intensity level

## TABLE 2.1

Number of storage bits for various values of $N$ and $k$.

| $\boldsymbol{N} / \boldsymbol{k}$ | $\mathbf{1}(\boldsymbol{L}=\mathbf{2})$ | $\mathbf{2}(\boldsymbol{L}=\mathbf{4})$ | $\mathbf{3}(\boldsymbol{L}=\mathbf{8})$ | $\mathbf{4}(\boldsymbol{L}=\mathbf{1 6})$ | $\mathbf{5}(\boldsymbol{L}=\mathbf{3 2})$ | $\mathbf{6}(\boldsymbol{L}=\mathbf{6 4})$ | $\mathbf{7}(\boldsymbol{L}=\mathbf{1 2 8})$ | $\mathbf{8}(\boldsymbol{L}=\mathbf{2 5 6})$ |
| ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| 32 | 1,024 | 2,048 | 3,072 | 4,096 | 5,120 | 6,144 | 7,168 | 8,192 |
| 64 | 4,096 | 8,192 | 12,288 | 16,384 | 20,480 | 24,576 | 28,672 | 32,768 |
| 128 | 16,384 | 32,768 | 49,152 | 65,536 | 81,920 | 98,304 | 114,688 | 131,072 |
| 256 | 65,536 | 131,072 | 196,608 | 262,144 | 327,680 | 393,216 | 458,752 | 524,288 |
| 512 | 262,144 | 524,288 | 786,432 | $1,048,576$ | $1,310,720$ | $1,572,864$ | $1,835,008$ | $2,097,152$ |
| 1024 | $1,048,576$ | $2,097,152$ | $3,145,728$ | $4,194,304$ | $5,242,880$ | $6,291,456$ | $7,340,032$ | $8,388,608$ |
| 2048 | $4,194,304$ | $8,388,608$ | $12,582,912$ | $16,777,216$ | $20,971,520$ | $25,165,824$ | $29,369,128$ | $33,554,432$ |
| 4096 | $16,777,216$ | $33,554,432$ | $50,331,648$ | $67,108,864$ | $83,886,080$ | $100,663,296$ | $117,440,512$ | $134,217,728$ |
| 8192 | $67,108,864$ | $134,217,728$ | $201,326,592$ | $268,435,456$ | $335,544,320$ | $402,653,184$ | $469,762,048$ | $536,870,912$ |

- Discrete intensity interval [0, L-1], $\mathrm{L}=2^{\mathrm{k}}$
- The number b of bits required to store a $\mathrm{M} \times \mathrm{N}$ digitized image
$\mathrm{b}=\mathrm{M} \times \mathrm{N} \times \mathrm{k}$


## Spatial and Intensity Resolution

- Spatial resolution
- A measure of the smallest discernible detail in an image
- stated with line pairs per unit distance, dots (pixels) per unit distance, dots per inch (dpi)
- Intensity resolution
- The smallest discernible change in intensity level
- stated with 8 bits, 12 bits, 16 bits, etc.

a b
c d
FIGURE 2.20 Typical effects of reducing spatial resolution. Images shown at: (a) 1250 dpi, (b) 300 dpi , (c) 150 dpi , and (d) 72 dpi . The thin black borders were added for clarity. They are not part of the data.

(Continued)
(e)-(h) Image (e)-(h) Image
displayed in 16,8, 4 , and 2 gray levels. (Original courtesy of
Dr. David
R. Pickens,
R. Pickens,
Department of

Radiology \&
Radiological
Sciences,
Sciences,
Vanderbilt
Vanderbilt
University
Medical Center.)


ab c
FIGURE 2.22 (a) Image with a low level of detail. (b) Image with a medium level of detail. (c) Image with a relatively large amount of detail. (Image (b) courtesy of the Massachusetts Institute of Technology.)


FIGURE 2.23
Typical
isopreference
curves for the three types of images in Fig. 2.22.

- Curves show isopreference points
- For low-, moderate-detail images, increasing K or $\mathbf{N}$ will help perceptual quality
- For high-detail images, increasing $\mathbf{N}$ is more important.


## Image Interpolation

- Interpolation - Process of using known data to estimate unknown values
e.g., zooming, shrinking, rotating, and geometric correction
- Interpolation (sometimes called resampling) - an imaging method to increase (or decrease) the number of pixels in a digital image.
Some digital cameras use interpolation to produce a larger image than the sensor captured or to create digital zoom
http://www.dpreview.com/learn/?/key=interpolation


## Image Interpolation:

Nearest Neighbor Interpolation


## Image Interpolation: <br> Bilinear Interpolation




## Image Interpolation:

## Bicubic Interpolation

- The intensity value assigned to point $(\mathrm{x}, \mathrm{y})$ is obtained by the following equation

$$
f_{3}(x, y)=\sum_{i=0}^{3} \sum_{j=0}^{3} a_{i j} x^{i} y^{j}
$$

- The sixteen coefficients are determined by using the sixteen nearest neighbors. http://en.wikipedia.org/wiki/Bicubic_interpolation

abc
def
FIGURE 2.24 (a) Image reduced to 72 dpi and zoomed back to its original size ( $3692 \times 2812$ pixels) using nearest neighbor interpolation. This figure is the same as Fig. 2.20(d). (b) Image shrunk and zoomed using bilinear interpolation. (c) Same as (b) but using bicubic interpolation. (d)-(f) Same sequence, but shrinking down to 150 dpi instead of 72 dpi [Fig. 2.24(d) is the same as Fig. 2.20(c)]. Compare Figs. 2.24(e) and (f), especially the latter, with the original image in Fig. 2.20(a).

| $\begin{array}{lll}0 & 1 & 1\end{array}$ |  | 0 | 1--1 |  | 0 |  | -1 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 010 |  | 0 | 1.0 |  | 0 |  | 0 |
| $0 \quad 0 \quad 1$ |  | 0 | $0^{`} 1$ |  | 0 |  |  |
| $\left.\begin{array}{lll}1 & 1 & 1\end{array}\right)$ | 0 | 0 | $0 \quad 0$ | 0 | 0 |  | 0 |
| $\left.1 \begin{array}{lll}1 & 0 & 1\end{array}\right\} R_{i}$ | 0 | 1 | 10 | 0 | 0 |  | 0 |
| 0 亿, 0 | 0 | 1 | 10 | 0 | 0 |  | 0 |
| 0 - ì | 0 |  | 11 1 | 0 | 0 |  | 0 |
| $\left.\begin{array}{lll}1 & 1 & 1\end{array}\right\} R_{j}$ | 0 | 1 | 11 | 0 | 0 |  | 0 |
| $\left.\begin{array}{lll}1 & 1 & 1\end{array}\right)^{\text {a }}$ | 0 | 0 | 00 | 0 | 0 |  |  |

a b c
$\begin{array}{lll}\mathrm{a} & \mathrm{b} & \mathrm{c} \\ \mathrm{d}\end{array}$
FIGURE 2.25 (a) An arrangement of pixels. (b) Pixels that are 8-adjacent (adjacency is shown by dashed lines; note the ambiguity). (c) $m$-adjacency. (d) Two regions that are adjacent if 8 -adjecency is used. (e) The circled point is part of the boundary of the 1 -valued pixels only if 8 -adjacency between the region and background is used. (f) The inner boundary of the 1-valued region does not form a closed path, but its outer boundary does.

Example: Addition of Noisy Images for Noise Reduction

Noiseless image: $f(x, y)$
Noise: $\mathrm{n}(\mathrm{x}, \mathrm{y})$ (at every pair of coordinates ( $\mathrm{x}, \mathrm{y}$ ), the noise is uncorrelated and has zero average value)
Corrupted image: $g(x, y)$

$$
g(x, y)=f(x, y)+n(x, y)
$$

Reducing the noise by adding a set of noisy images, $\left\{\mathrm{g}_{\mathrm{i}}(\mathrm{x}, \mathrm{y})\right\}$

$$
\bar{g}(x, y)=\frac{1}{K} \sum_{i=1}^{K} g_{i}(x, y)
$$

Example: Addition of Noisy Images for Noise Reduction

$$
\begin{aligned}
& \quad \bar{g}(x, y)=\frac{1}{K} \sum_{i=1}^{K} g_{i}(x, y) \\
& E\{\bar{g}(x, y)\}=E\left\{\frac{1}{K} \sum_{i=1}^{K} g_{i}(x, y)\right\} \quad \sigma_{\bar{g}(x, y}^{2} \\
& =E\left\{\frac{1}{K} \sum_{i=1}^{K}\left[f(x, y)+n_{i}(x, y)\right]\right\} \\
& = \\
& =f(x, y)+E\left\{\frac{1}{K} \sum_{i=1}^{K} n_{i}(x, y)\right\} \\
& =
\end{aligned} \quad=\sigma^{2}(x, y) \quad l
$$

Example: Addition of Noisy Images for Noise Reduction

In astronomy, imaging under very low light levels frequently causes sensor noise to render single images virtually useless for analysis.

- In astronomical observations, similar sensors for noise reduction by observing the same scene over long periods of time. Image averaging is then used to reduce the noise.

a b c
d e f
FIGURE 2.26 (a) Image of Galaxy Pair NGC 3314 corrupted by additive Gaussian noise. (b)-(f) Results of averaging $5,10,20,50$, and 100 noisy images, respectively. (Original image courtesy of NASA.)



## Black regions in (c) shows no difference Non black is regions where they are different

## An Example of Image Subtraction: Mask Mode Radiography

Mask h(x,y): an X-ray image of a region of a patient's body

Live images $\mathbf{f}(\mathbf{x}, \mathbf{y})$ : X-ray images captured at TV rates after injection of the contrast medium

Enhanced detail g(x,y)

$$
g(x, y)=f(x, y)-h(x, y)
$$


a b
c d
FIGURE 2.28
Digital subtraction angiography. (a) Mask image. (b) A live image. (c) Difference between (a) and (b). (d) Enhanced difference image. (Figures (a) and (b) courtesy of

The Image
Sciences Institute, University Medical Center, Utrecht, The Netherlands.)

The procedure gives a movie showing how the contrast medium propagates through the various arteries in the area being observed.

## An Example of Image Multiplication


a b c
FIGURE 2.29 Shading correction. (a) Shaded SEM image of a tungsten filament and support, magnified approximately 130 times. (b) The shading pattern. (c) Product of (a) by the reciprocal of (b). (Original image courtesy of Mr. Michael Shaffer, Department of Geological Sciences, University of Oregon, Eugene.)

a b c
FIGURE 2.30 (a) Digital dental X-ray image. (b) ROI mask for isolating teeth with fillings (white corresponds to 1 and black corresponds to 0). (c) Product of (a) and (b).


## a b c <br> d e

FIGURE 2.31
(a) Two sets of coordinates, $A$ and $B$, in 2-D space. (b) The union of $A$ and $B$.

(c) The intersection of $A$ and $B$. (d) The complement of $A$. (e) The difference between $A$ and $B$. In (b)-(e) the shaded areas represent the member of the set operation indicated.

a b c
FIGURE 2.32 Set
operations
involving gray-
scale images.
(a) Original
image. (b) Image
negative obtained
using set
complementation.
(c) The union of
(a) and a constant
image.
(Original image
courtesy of G.E.


FIGURE 2.33
Illustration of logical operations involving foreground (white) pixels. Black represents binary 0s and white binary 1s. The dashed lines are shown for reference only. They are not part of the result.


FIGURE 2.34 Intensity transformation function used to obtain the negative of an 8 -bit image. The dashed arrows show transformation of an arbitrary input intensity value $z_{0}$ into its corresponding output value $S_{0}$.
Imaner

a b
c d
FIGURE 2.35
Local averaging using neighborhood processing. The procedure is illustrated in (a) and (b) for a rectangular neighborhood. (c) The aortic angiogram discussed in Section 1.3.2. (d) The result of using Eq. (2.6-21) with $m=n=41$. The images are of size $790 \times 686$ pixels.

## Geometric Spatial Transformations

- Geometric transformation (rubber-sheet transformation)
- A spatial transformation of coordinates

$$
(x, y)=T\{(v, w)\}
$$

— intensity interpolation that assigns intensity values to the spatially transformed pixels.

- Affine transform

$$
\left[\begin{array}{lll}
x & y & 1
\end{array}\right]=\left[\begin{array}{lll}
v & w & 1
\end{array}\right]\left[\begin{array}{lll}
t_{11} & t_{12} & 0 \\
t_{21} & t_{22} & 0 \\
t_{31} & t_{32} & 1
\end{array}\right]
$$

TABLE 2.2
Affine transformations based on Eq. (2.6.-23).

| Transformation Name | Affine Matrix, $T$ | Coordinate Equations | Example |
| :---: | :---: | :---: | :---: |
| Identity | $\left[\begin{array}{lll}1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1\end{array}\right]$ | $x=v$ $y=w$ |  |
| Scaling | $\left[\begin{array}{lll}c_{x} & 0 & 0 \\ 0 & c_{y} & 0 \\ 0 & 0 & 1\end{array}\right]$ | $x=c_{x} v$ $y=c_{y} w$ |  |
| Rotation | $\left[\begin{array}{ccc}\cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1\end{array}\right]$ | $\begin{aligned} & x=v \cos \theta-w \sin \theta \\ & y=v \cos \theta+w \sin \theta \end{aligned}$ |  |
| Translation | $\left[\begin{array}{ccc}1 & 0 & 0 \\ 0 & 1 & 0 \\ t_{x} & t_{y} & 1\end{array}\right]$ | $x=v+t_{x}$ $y=w+t_{y}$ |  |
| Shear (vertical) | $\left[\begin{array}{lll}1 & 0 & 0 \\ s_{v} & 1 & 0 \\ 0 & 0 & 1\end{array}\right]$ | $\begin{gathered} x=v+s_{v} w \\ y=w \end{gathered}$ |  |
| Shear (horizontal) | $\left[\begin{array}{ccc}1 & s_{h} & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1\end{array}\right]$ | $\begin{gathered} x=v \\ y=s_{h} v+w \end{gathered}$ |  |



FIGURE 2.36 (a) A 300 dpi image of the letter T. (b) Image rotated $21^{\circ}$ clockwise using nearest neighbor interpolation to assign intensity values to the spatially transformed pixels. (c) Image rotated $21^{\circ}$ using bilinear interpolation. (d) Image rotated $21^{\circ}$ using bicubic interpolation. The enlarged sections show edge detail for the three interpolation approaches.

## Image Registration

- Input and output images are available but the transformation function is unknown.

Goal: estimate the transformation function and use it to register the two images.

- One of the principal approaches for image registration is to use tie points (also called control points)
$>$ The corresponding points are known precisely in the input and output (reference) images.


## Image Registration

- A simple model based on bilinear approximation:

$$
\left\{\begin{array}{l}
x=c_{1} v+c_{2} w+c_{3} v w+c_{4} \\
y=c_{5} v+c_{6} w+c_{7} v w+c_{8}
\end{array}\right.
$$

Where $(v, w)$ and $(x, y)$ are the coordinates of tie points in the input and reference images.

a b
c d
FIGURE 2.37
Image
registration. (a) Reference image. (b) Input (geometrically distorted image) Corresponding tie points are shown as small white squares near the corners.
(c) Registered image (note the errors in the borders). (d) Difference between (a) and (c), showing more registration errors.


## Image Transform

- A particularly important class of 2-D linear transforms, denoted $T(u, v)$
$T(u, v)=\sum_{x=0}^{M-1} \sum_{y=0}^{N-1} f(x, y) r(x, y, u, v)$
where $f(x, y)$ is the input image,
$r(x, y, u, v)$ is the forward transformation ker nel, variables $u$ and $v$ are the transform variables,

$$
u=0,1,2, \ldots, \mathrm{M}-1 \text { and } v=0,1, \ldots, \mathrm{~N}-1 .
$$

## Image Transform

- Given $T(u, v)$, the original image $f(x, y)$ can be recoverd using the inverse tranformation of $\mathrm{T}(\mathrm{u}, \mathrm{v})$.
$f(x, y)=\sum_{u=0}^{M-1} \sum_{v=0}^{N-1} T(u, v) s(x, y, u, v)$
where $s(x, y, u, v)$ is the inverse transformation ker nel,
$x=0,1,2, \ldots, \mathrm{M}-1$ and $y=0,1, \ldots, \mathrm{~N}-1$.



| $a$ | $b$ |
| :--- | :--- |
| $c$ | $d$ |

FIGURE 2.40
(a) Image corrupted
by sinusoidal
interference. (b)
Magnitude of the
Fourier transform
showing the bursts
of energy responsible for the interference.
(c) Mask used to
eliminate the energy
bursts. (d) Result of
computing the inverse of the modified Fourier transform. (Original image courtesy of NASA.)

## Forward Transform Kernel

$T(u, v)=\sum_{x=0}^{M-1} \sum_{y=0}^{N-1} f(x, y) r(x, y, u, v)$

The kernel $r(x, y, u, v)$ is said to be SEPERABLE if $r(x, y, u, v)=r_{1}(x, u) r_{2}(y, v)$

In addition, the kernel is said to be SYMMETRIC if $r_{1}(x, u)$ is functionally equal to $r_{2}(y, v)$, so that $r(x, y, u, v)=r_{1}(x, u) r_{1}(y, u)$

## The Kernels for 2-D Fourier Transform

The forward kernel
$r(x, y, u, v)=e^{-j 2 \pi(u x / M+v y / N)}$
Where $j=\sqrt{-1}$

The inverse kernel
$s(x, y, u, v)=\frac{1}{M N} e^{j 2 \pi(u x / M+v y / N)}$

## 2-D Fourier Transform

$$
\begin{aligned}
& T(u, v)=\sum_{x=0}^{M-1} \sum_{y=0}^{N-1} f(x, y) e^{-j 2 \pi(u x / M+v y / N)} \\
& f(x, y)=\frac{1}{M N} \sum_{u=0}^{M-1} \sum_{v=0}^{N-1} T(u, v) e^{j 2 \pi(u x / M+v y / N)}
\end{aligned}
$$

## Probabilistic Methods

Let $z_{i}, i=0,1,2, \ldots, L-1$, denote the values of all possible intensities in an $M \times N$ digital image. The probability, $p\left(z_{k}\right)$, of intensity level
$z_{k}$ occurring in a given image is estimated as

$$
p\left(z_{k}\right)=\frac{n_{k}}{M N},
$$

where $n_{k}$ is the number of times that intensity $z_{k}$ occurs in the image.

$$
\sum_{k=0}^{L-1} p\left(z_{k}\right)=1
$$

The mean (average) intensity is given by

$$
m=\sum_{k=0}^{L-1} z_{k} p\left(z_{k}\right)
$$

## Probabilistic Methods

The variance of the intensities is given by

$$
\sigma^{2}=\sum_{k=0}^{L-1}\left(z_{k}-m\right)^{2} p\left(z_{k}\right)
$$

The $n^{\text {th }}$ moment of the intensity variable $z$ is

$$
u_{n}(z)=\sum_{k=0}^{L-1}\left(z_{k}-m\right)^{n} p\left(z_{k}\right)
$$


a b c
FIGURE 2.41
Images exhibiting
(a) low contrast,
(b) medium
contrast, and
(c) high contrast.
$\sigma=14.3$

